

The Latin American Giant Observatory (LAGO) Project

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- 3 Data Acquisition System
- 4 Preliminary Results
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The Latin American Giant Observatory

The LAGO Collaboration

The LAGO Collaboration

- 99 members from 30 institutions in 11 countries: Argentina, Bolivia, Brasil, Chile, Colombia, Ecuador, Guatemala, México, Perú, Venezuela and Spain.
- LAGO is developing a network of cosmic ray detectors based on Water Cherenkov Detectors
- LAGO detectors span 90° latitude (from Mexico to Antarctica) and 5230 meters of altitude (from 10 to 5240 masl).



The Latin American Giant Observatory

Goals & Activities

Scientific goals

- Astroparticle physics to study the Extreme Universe (EU)
- Transient and long term Space Weather phenomena through Solar modulation of Cosmic Rays (CR)
- Measurements of background radiation at ground level



Academic goals

- To Train Latin American students in HE and Astroparticle physics
- To Build a LA network of Astroparticle researchers

The Latin American Giant Observatory

Organization

How does it works?

- Non-centralized, collaborative network of institutions and scientists
- 3 working groups, 9+2 members coordination committee, 1 P.I.
- Developments, expertise and data are shared across the network.
- Scientific goals are developed by dedicated LAGO working groups.
- Academic goals are developed trough Network schools and internships.

Working groups

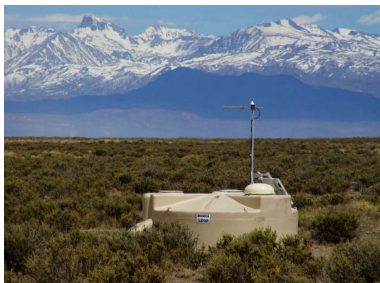
- Physics (GRBs & Space Weather)
- Data (Simulation & Analysis)
- Detectors (Design & Construction)

The Latin American Giant Observatory

Working groups

Origins

- LAGO project was born as "spin-off" of Pierre Auger Observatory Mendoza (ARG).
- Researchers of LAGO were trained at PAO.
- First PMTs and electronics were recycled from PAO.



Scope of the Talk

In this talk we are going to describe the development of a low-budget WCD for LAGO project and future perspectives.

Water Cherenkov Detector

Block Diagram

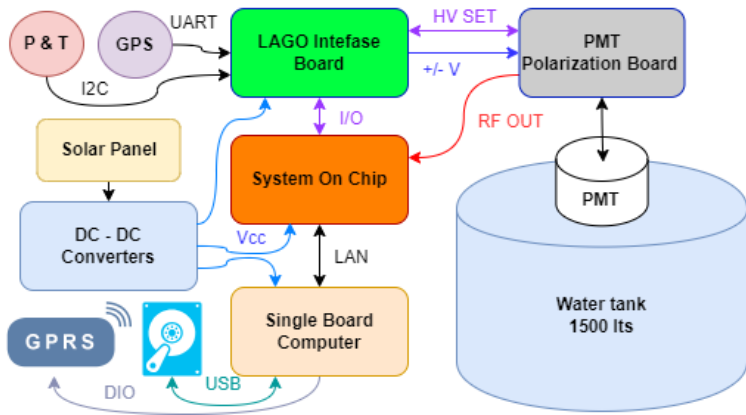


Figure: Block diagram of an autonomus LAGO WCD

Water Cherenkov Detector

Water tank

Characteristics

- Commercial PVC water tanks are commonly used (1 to 2.5 sqm).
- External coating to reflect sunlight is mandatory.
- Internal coating with high diffusive and reflective material is used (Tyvek®), $r = 0.97$ at $400nm$.
- Tank detection area $A_{Tank} = 1.58m^2$
- Ultra pure water is used after filtering and adding H_2O_2 (10ppm) to ensure transparency for a long period of time, $t = 0.99$ at $400nm$.



Water Cherenkov Detector

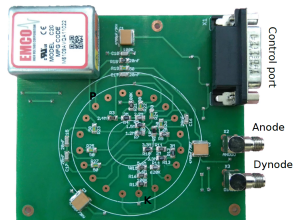
Photomultiplier

Characteristics

- 8" Photomultiplier 10 stages (Hamamatsu R5912).
- Quantum Efficiency $QE = 25\%$ at $400nm$.
- Detection area $A_{PMT} = 380cm^2$.
- Anode and last dynode signals used to enhance dynamic range.
- LAGO designed a board to power, control and signal conditioning.
- Isolating PMT polarization board from moisture and condensation is crucial.



EMCO C20



Water Cherenkov Detector

Water tank

Photo-electrons produced inside the tank

$$N_{pe} = N_{phot/cm} \times H \times QE \times \left(\frac{A_{PMT}}{A_{Tank}} \right) \times r \times t$$

Where:

- $N_{phot/cm}$ = amount of Cherenkov photons per cm [320].
- H = water height in cm.
- QE = PMT Quantum Efficiency.
- A_{PMT}/A_{Tank} = ratio between PMT detection area and Tank detection area.
- r = internal coating reflection coefficient
- t = water transmission coefficient.

For the WCD presented here 120 photons/cm are expected

Characteristics

- 2 fast inputs ($\pm 1V$ $1M\Omega$ digitized at $125Msps$, 14 bit).
- 4 slow inputs ($0 - 3.5V$ digitized at $100ksps$, 12 bit).
- 4 analog outputs ($(0 - 1.8V, PWM, 12$ bit).
- Xlink System On Chip (Dual-Core ARM Cortex-A9 μP + FPGA Xilinx Zynq 7010).
- Ethernet 1GB, Wifi, USB 2.0.
- I2C, UART, SPI.



STEM LAB 125-14
cost under 500 EUR.

Characteristics

- +12V 1A input.
- +12V, +5V, $\pm 3.3V$ outputs.
- PMT HV source output control (0-2.5V ramp).
- PMT HV sense (0 - 3.3V).
- PMT temperature sensor reading.
- Pulse Per Second signal form GPS for synchronization.
- Atmospheric Pressure and Temperature monitoring.



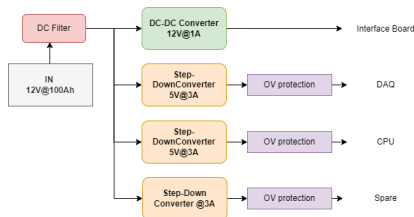
Leopard Interface board
cost under 200 EUR.

Data Acquisition System

Power Distribution Board

Characteristics

- +12V at 2A input.
- +12V limited 1A.
- 3 +5V at 3A outputs.
- Over-current protection.
- Isolate Solar panel battery from DAQ system.



- Designed with commercial modules used for industrial applications
- Meet safety and compliance standards.
- cost under 100 EUR.

Data Acquisition System

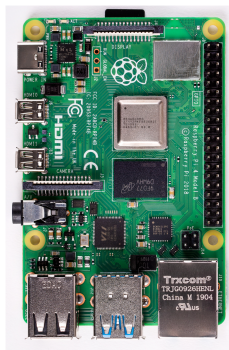
Data Storage and Communication

Description

A Single Board Computer (Raspberry Pi 4) is used for configuration, data storage and communication.

Characteristics

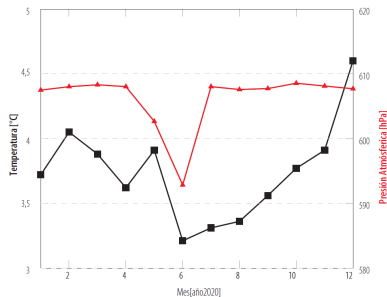
- DAQ configuration scripts (PMT setup, trigger, metadata)
- Every hour a text file is saved.
- Wifi /Ethernet data transfer if available.
- GPRS communication (only for system status)
- 1TB Solid State Disk for local storage.



Preliminary studies

Location

- 2 WCDs installed near Chimborazo volcano at 4330 m.a.s.l.
- Stainless steel tanks will be deployed.

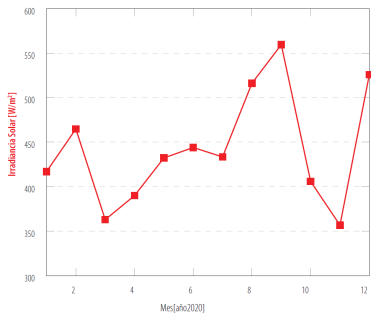
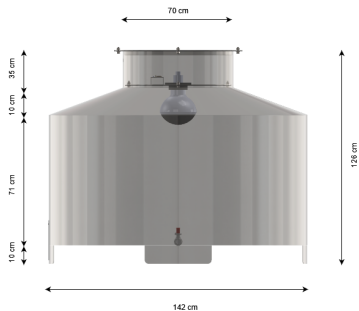


Atmospheric pressure: Max = 608hPa, Min = 598hPa

Temperature: Max = 4.6°C, Min = 3.2°C

Power generation & Communication

- DAQ powered with a 12V 100Ah solar panel.
- Only GPRS signal is available.

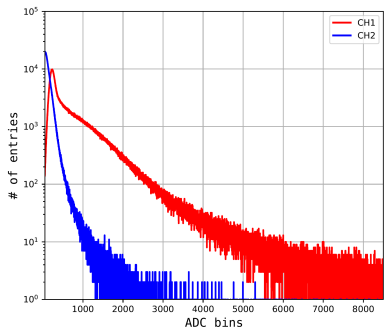
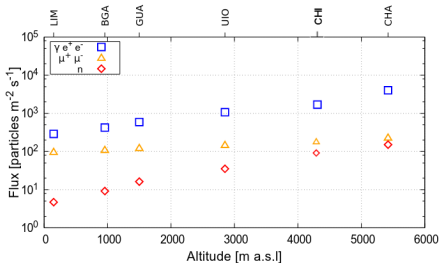


Medium Irradiance expected 450W/m² equivalent to 390W at solar panel

Preliminary studies

Particle flux & Signal

- Expected flux simulated with ARTI framework.
- Preliminary data analyzed.

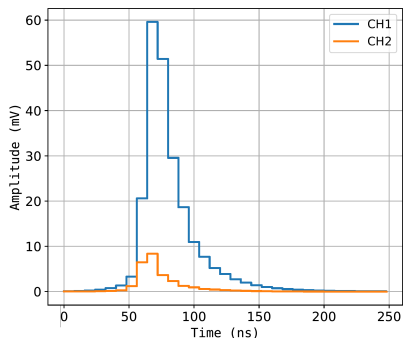
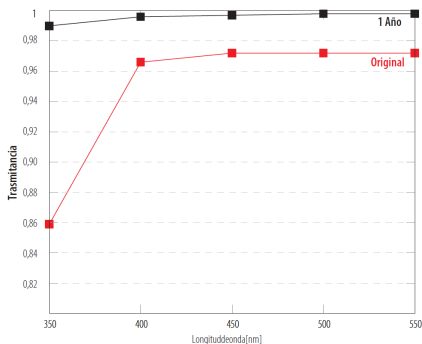


Signal rate expected 2KHz. PMT HV level and trigger must be fine tuned.

Preliminary studies

Water transparency and Maintenance

- H_2O_2 shock treatment improves water transparency over time
- Time between water change minimum 2 years.

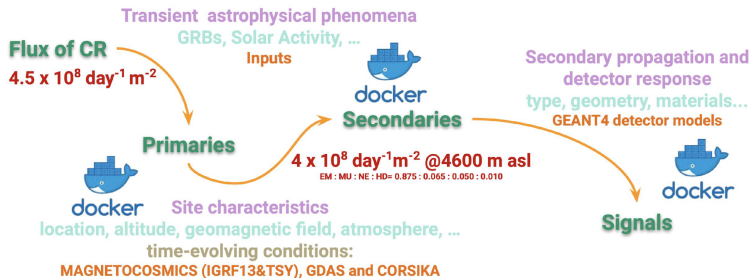


Signal level must be improved.

Implementation

- A simulation framework from Primary CR to WCD response ARTI available.
- ARTI is a blend of CORSIKA, MAGCOS and GEANT4

Our secondary flux simulation framework



So far

- An autonomous WCD detector can be implemented at low-budget (5K EUR),
- A small array of WCDs located at 4330 m.a.s.l. will be ready by March 2024.

The future

- WCD's performance at low energies could be improved adding a different CR detector.
- **RPC** detectors specially designed to work at high altitude are excellent candidates.

Further information can be found in www.lagoproject.net

grazie!

